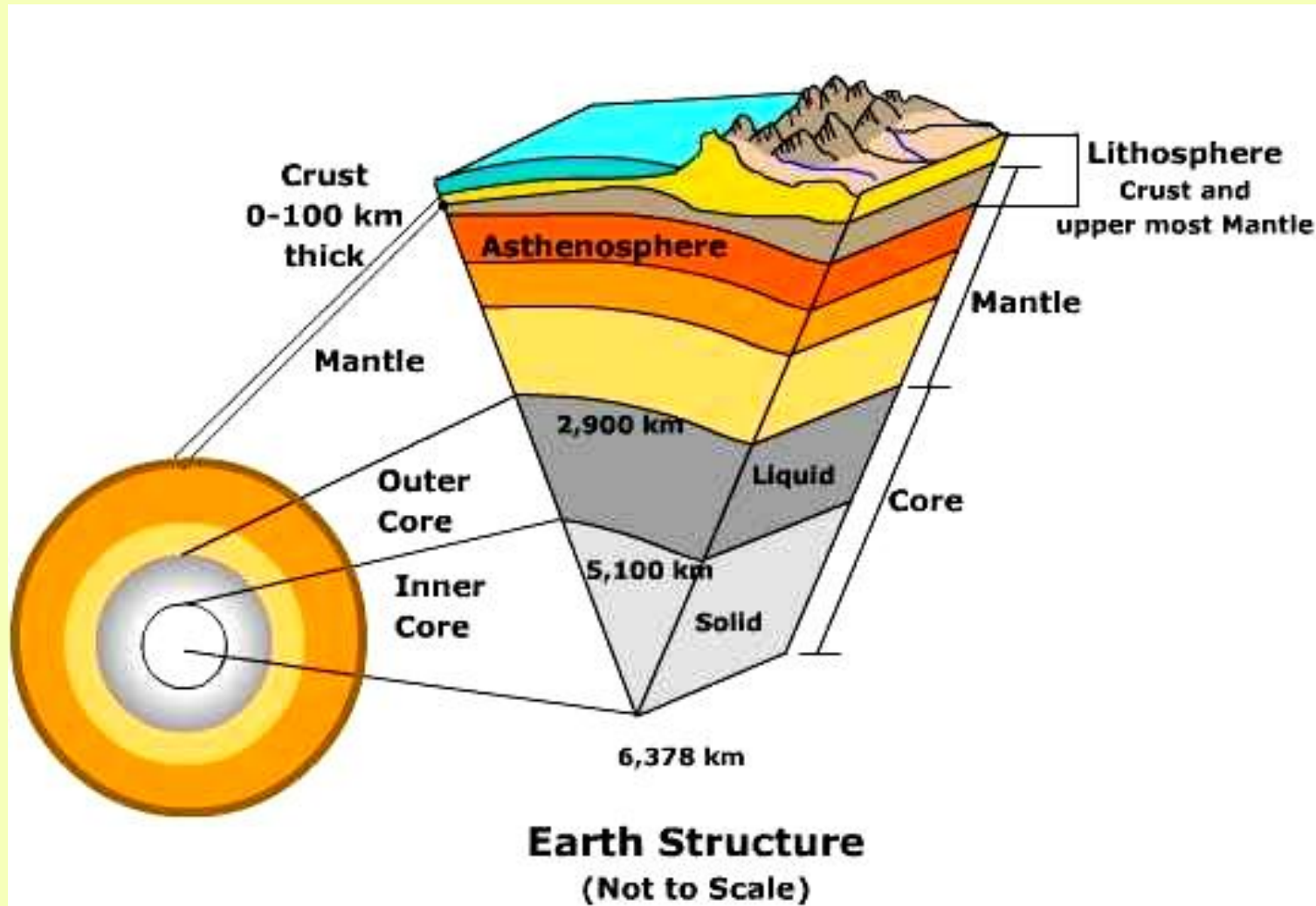
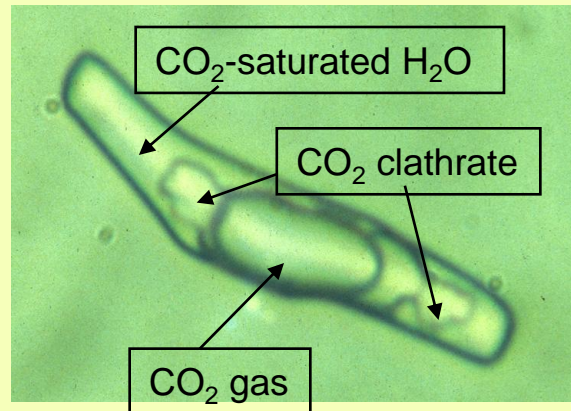
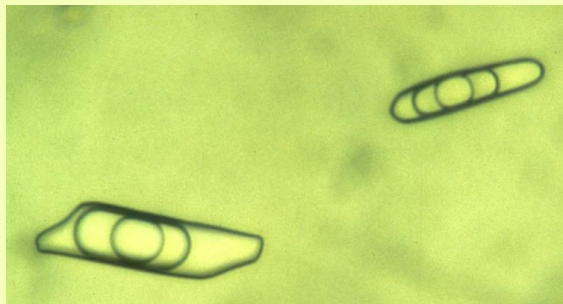
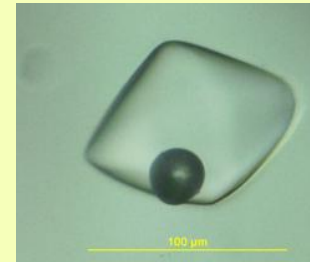
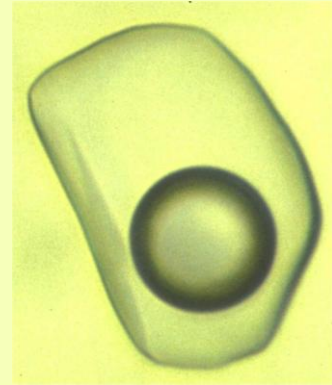


# Fluid and Melt Inclusion Evidence for the Distribution and Speciation of Carbon in the Earth

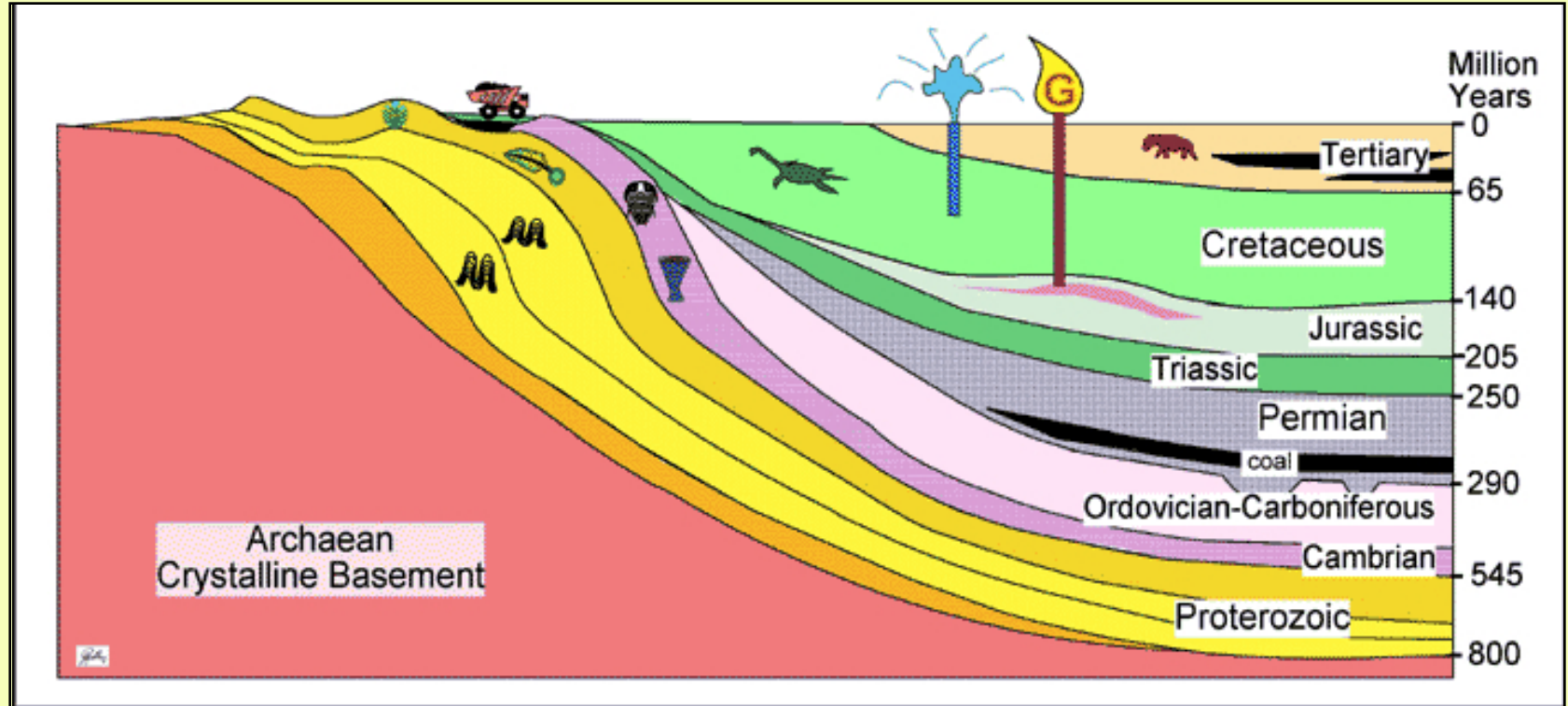


# Fluid and melt inclusions provide the most direct evidence for the nature of fluids in the lithosphere.

- *FI in sedimentary (diagenetic) environments*
- *FI in metamorphic environments*
- *FI & MI in shallow crustal magmatic environments*
- *FI & MI in upper mantle environments*
- *FI in diamonds*
- *FI in deeper mantle minerals (nano- to pico-scale FI)*

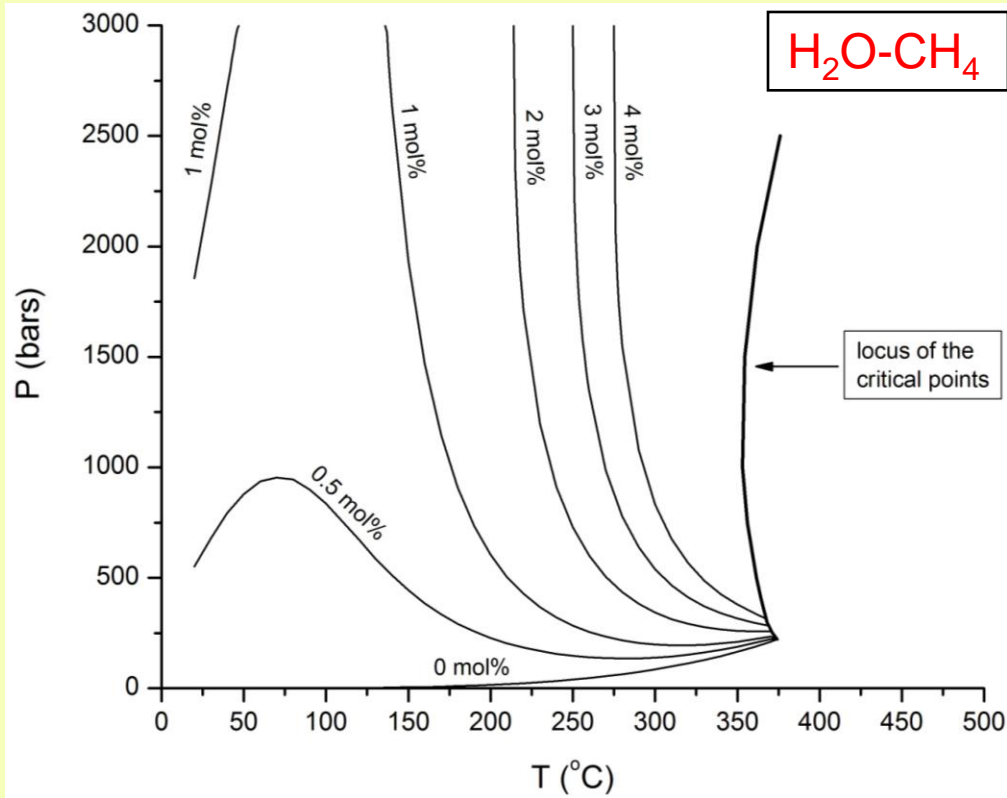


During the burial and diagenesis of organic-rich sediments the organic matter undergoes systematic changes with increasing depth of burial (temperature).

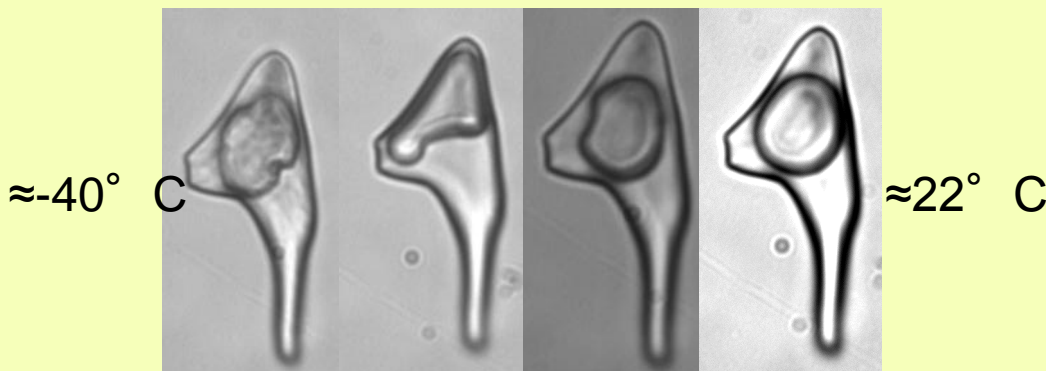
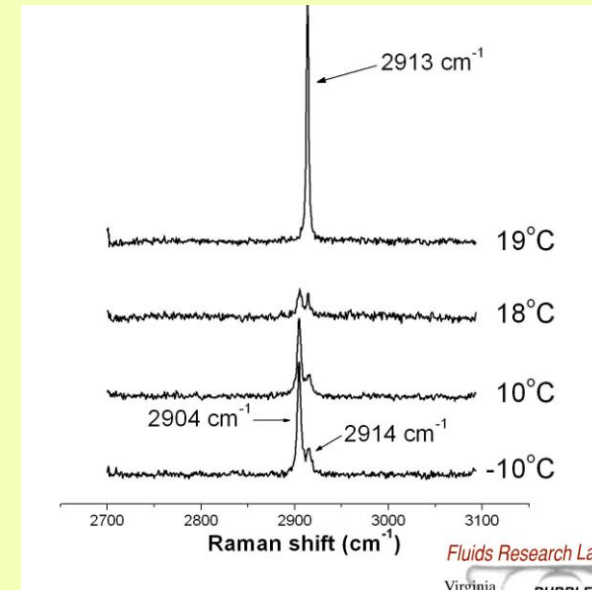


At the shallowest depths biogenic  $\text{CH}_4$  is produced as organic material starts to breakdown. The methane dissolves in pore waters which can be trapped to produce  $\text{CH}_4$ -bearing aqueous fluid inclusions. In submarine areas near continental margins and in permafrost regions the  $\text{CH}_4$  can react with  $\text{H}_2\text{O}$  to produce methane hydrate (clathrate)

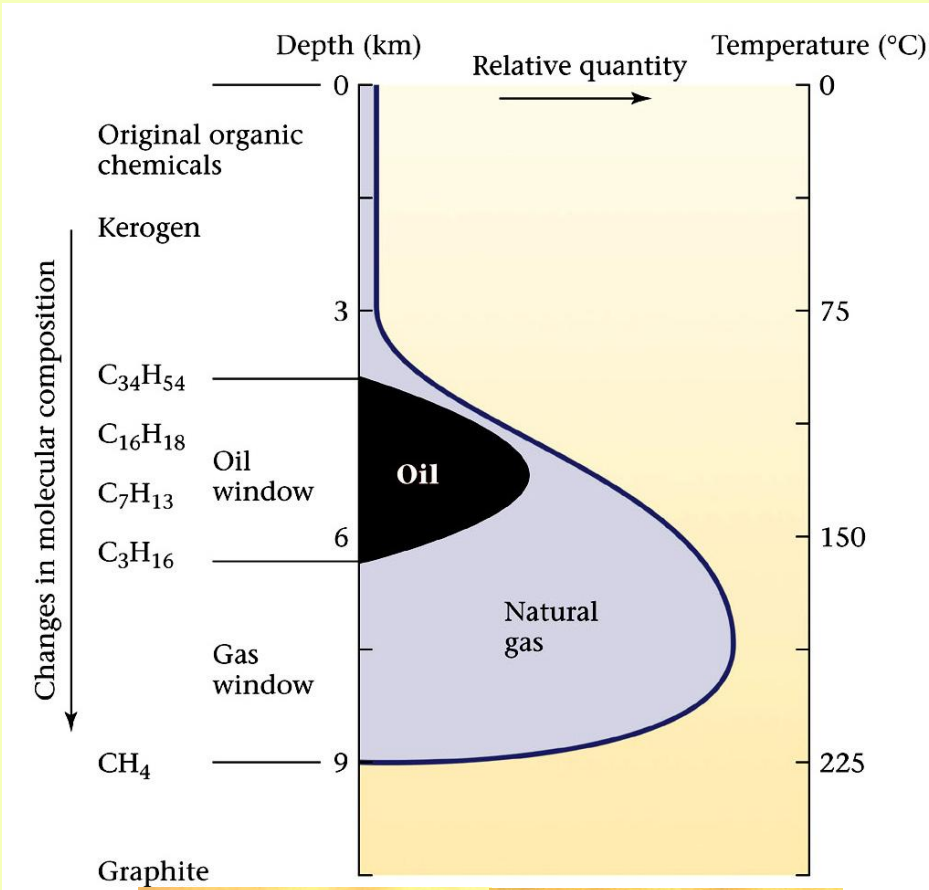
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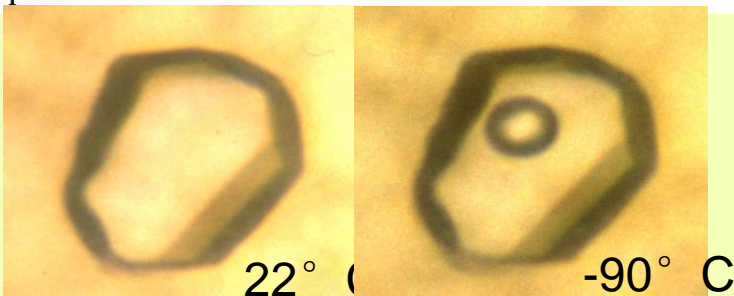
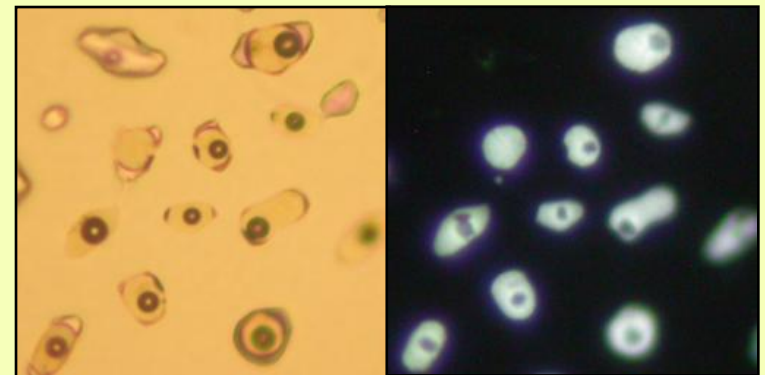
At low temperature ( $\approx < 30^\circ \text{C}$ ) and high pressure, methane hydrate (clathrate) is stable and occurs naturally in the sub-seafloor near continental margins and in permafrost regions. It is estimated that worldwide more methane exists in hydrate deposits ( $\approx 10^{17} - 10^{20}$  cubic feet!) than in all the other fossil energy (coal, oil, natural gas) deposits combined.



During the burial and diagenesis of organic-rich sediments the organic matter undergoes systematic changes with increasing depth of burial (temperature).

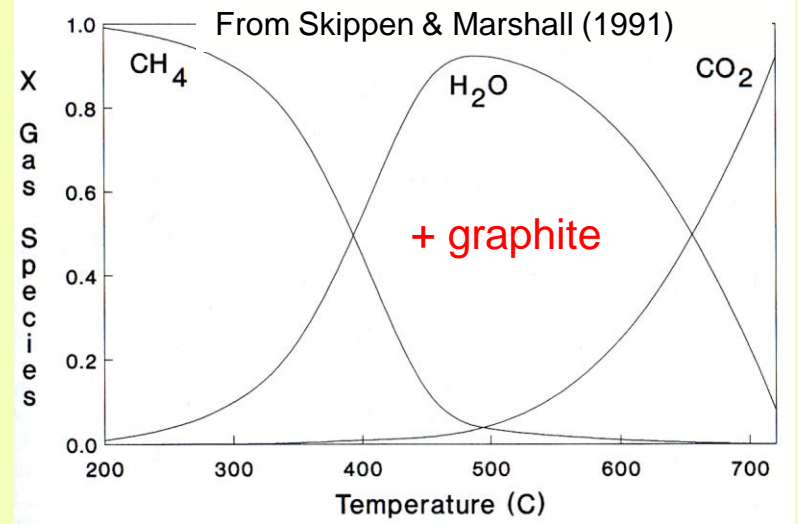
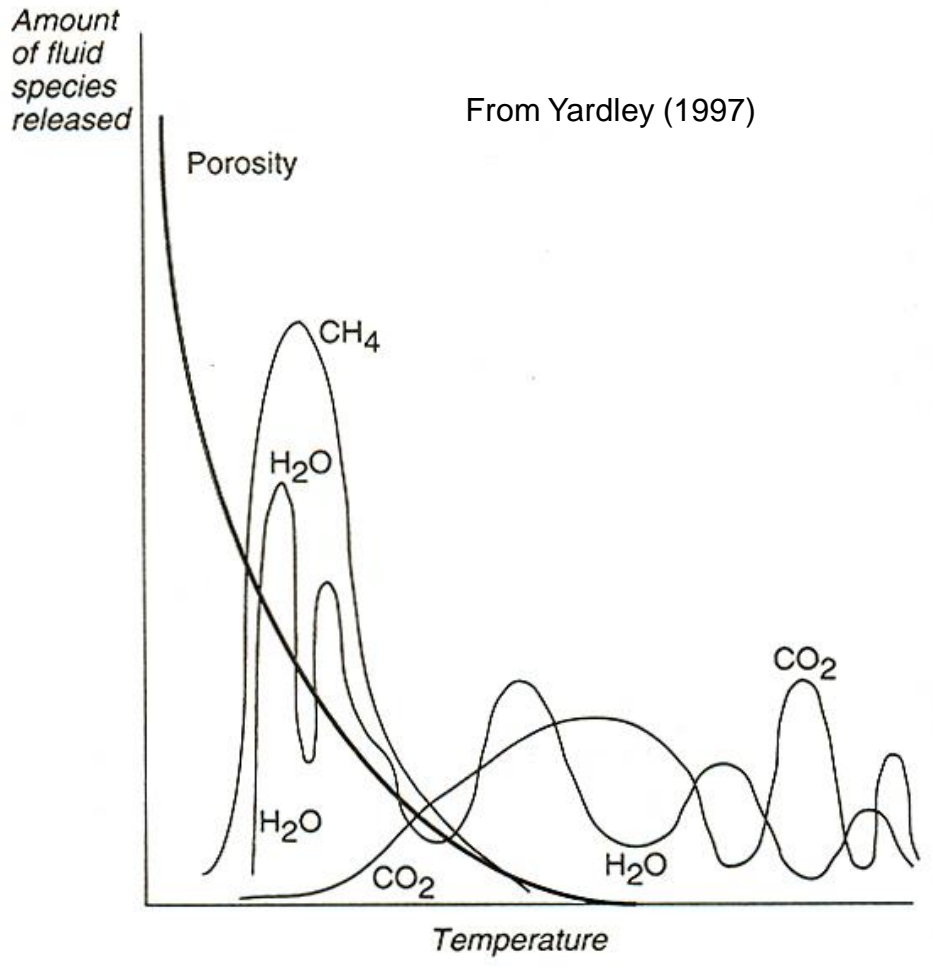


With continued burial the sediments enter the “oil window” where liquid hydrocarbons are generated. When the sediments are buried to depths greater than the “oil window” and enter the anchimetamorphic zone, liquid hydrocarbons become unstable and break down to produce mostly methane gas, with lesser amounts of ethane, propane, butane and non-hydrocarbon gases ( $N_2$ ,  $H_2S$ , etc.).



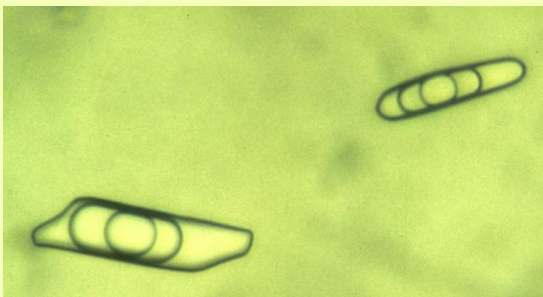
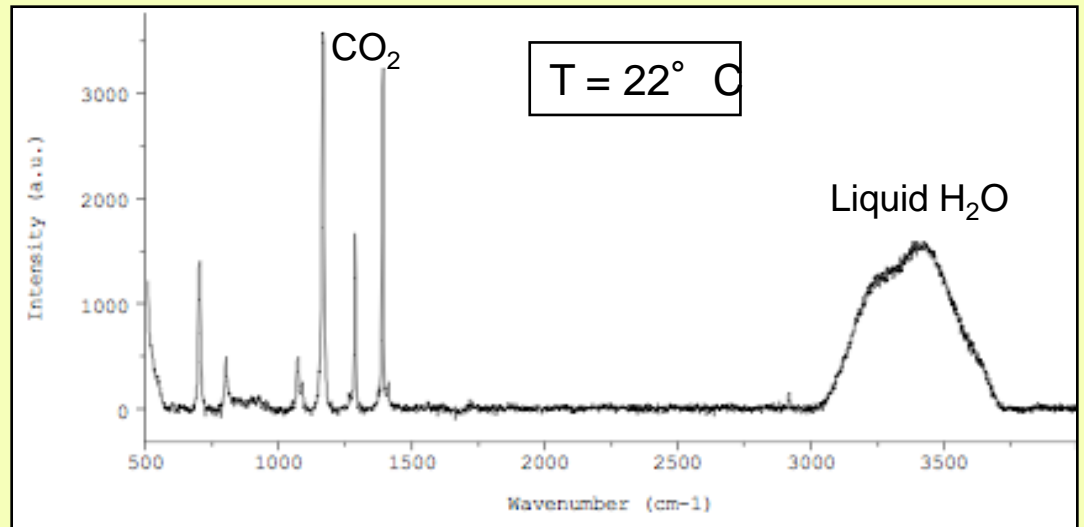
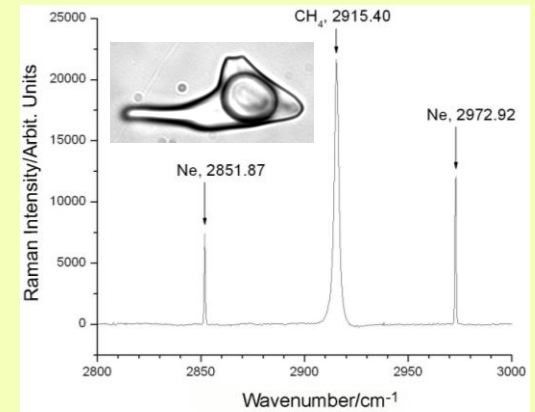
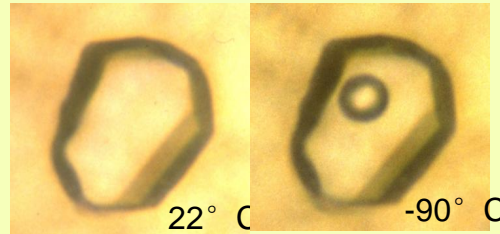
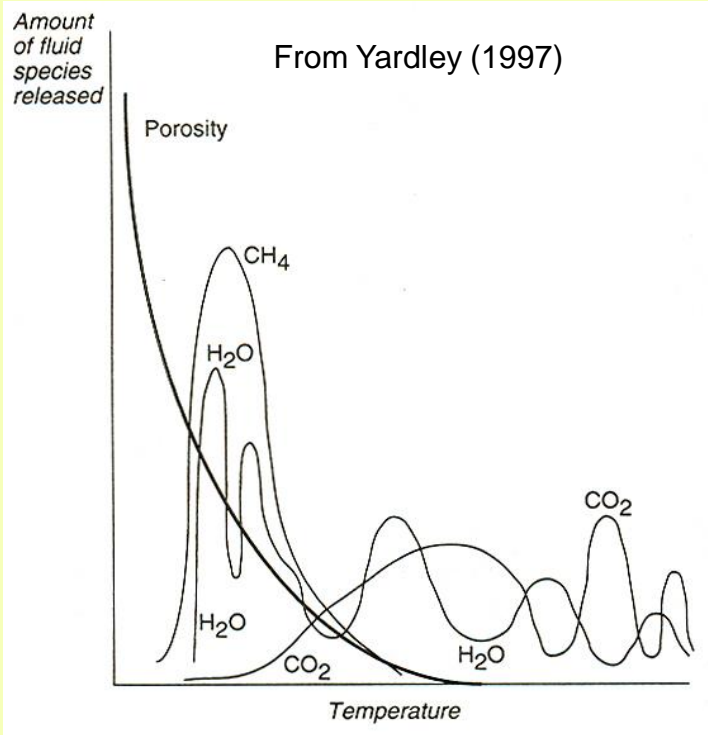
CARBON IN THE EARTH

Most metamorphic fluids are approximated by the C-H-O-S-(salt) system. The composition of metamorphic fluids varies systematically with changing  $f_{O_2}$ , temperature and pressure.



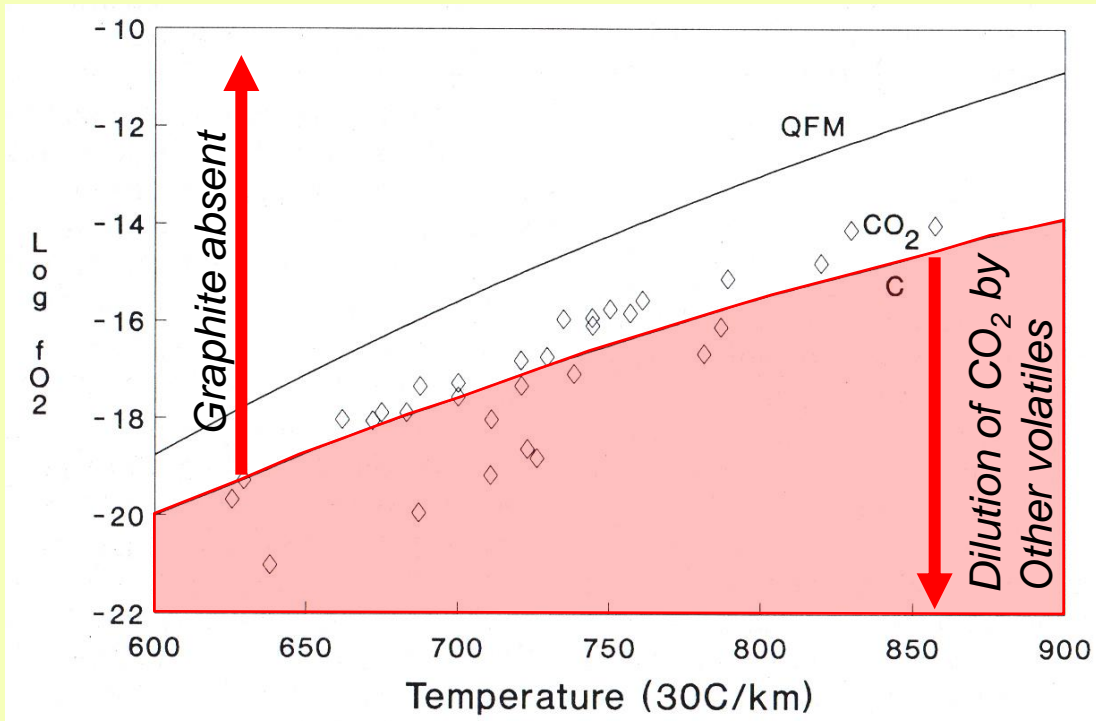
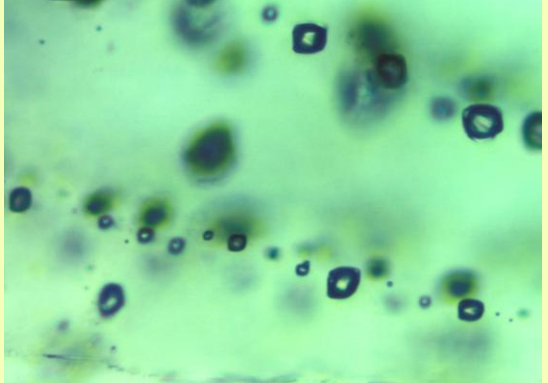
The relative proportions of fluid species vs. T for QFM-1 is shown above. As noted by Yardley (1997), the actual proportions of different fluid species will be controlled by fluid-mineral reactions and will involve pulses of CH<sub>4</sub>, H<sub>2</sub>O or CO<sub>2</sub>-rich fluids as organic matter, hydrous phases and carbonate minerals, respectively, break down at different temperatures.

Most metamorphic fluids are approximated by the C-H-O-S-(salt) system. The composition of metamorphic fluids varies systematically with changing  $f_{O_2}$ , temperature and pressure.



CARBON IN THE EARTH

Mineral equilibria for granulite grade metamorphic rocks indicate that the oxygen fugacity of the lower continental crust is  $\approx$ QFM - 1.



QFM - 1 for lower crustal rocks is consistent with the fact that most granulites are at (or close to) saturation in graphite and that nearly pure  $CO_2$  fluid inclusions are ubiquitous in granulites.

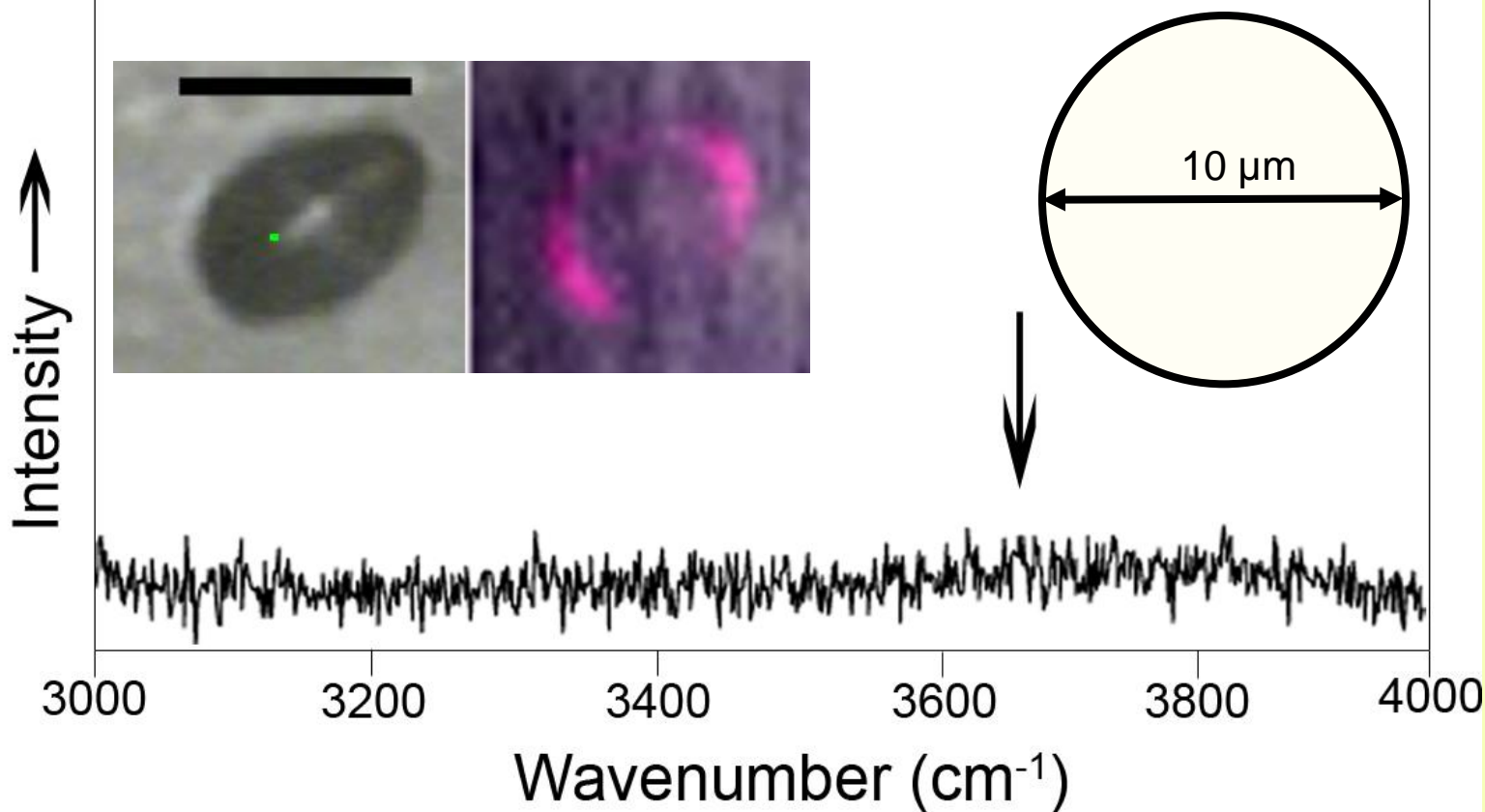
From Skippen & Marshall, 1991, Canadian Mineralogist

Diamonds above represent oxygen fugacities based on oxide barometry from Lamb & Valley (1984) and assuming a 30° C/km geotherm.

CARBON IN THE EARTH

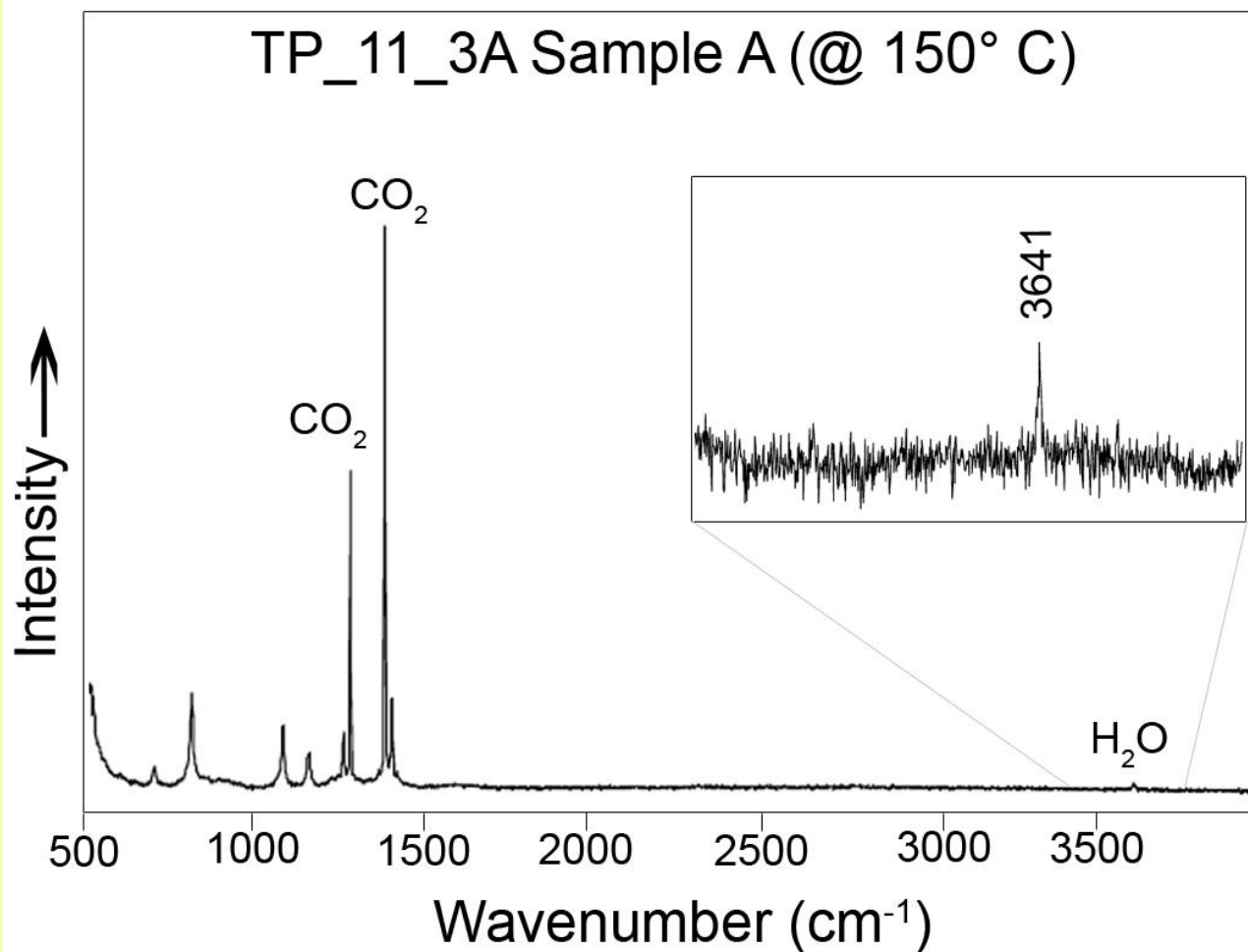
H<sub>2</sub>O is not detected during room temperature Raman analysis of “single” phase fluid inclusions in granulites from the Adirondack Mtns.

TP\_11\_3A Sample A (@ 27° C)



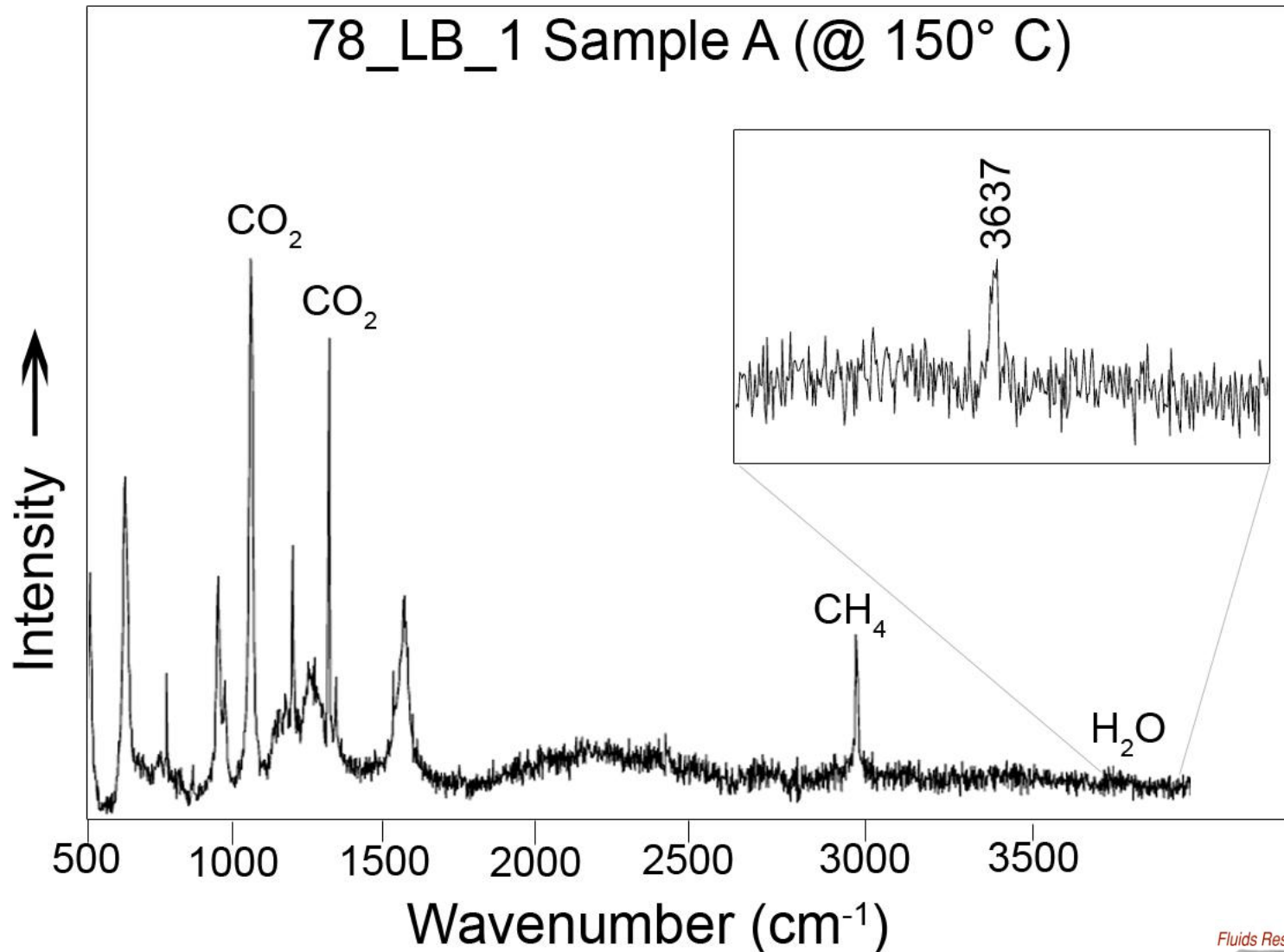
If a 10 μm spherical FI contains 90 mol% CO<sub>2</sub> with a density of 0.7 g/cc, and 10 mol% H<sub>2</sub>O with a density of 1 g/cc, the H<sub>2</sub>O phase will occur as a 0.05 μm (50 nanometer) film on the FI walls!

When the same FI is analyzed at 150° C, a small but distinct peak characteristic of H<sub>2</sub>O is observed. Results suggest a few mole% H<sub>2</sub>O in the inclusion.

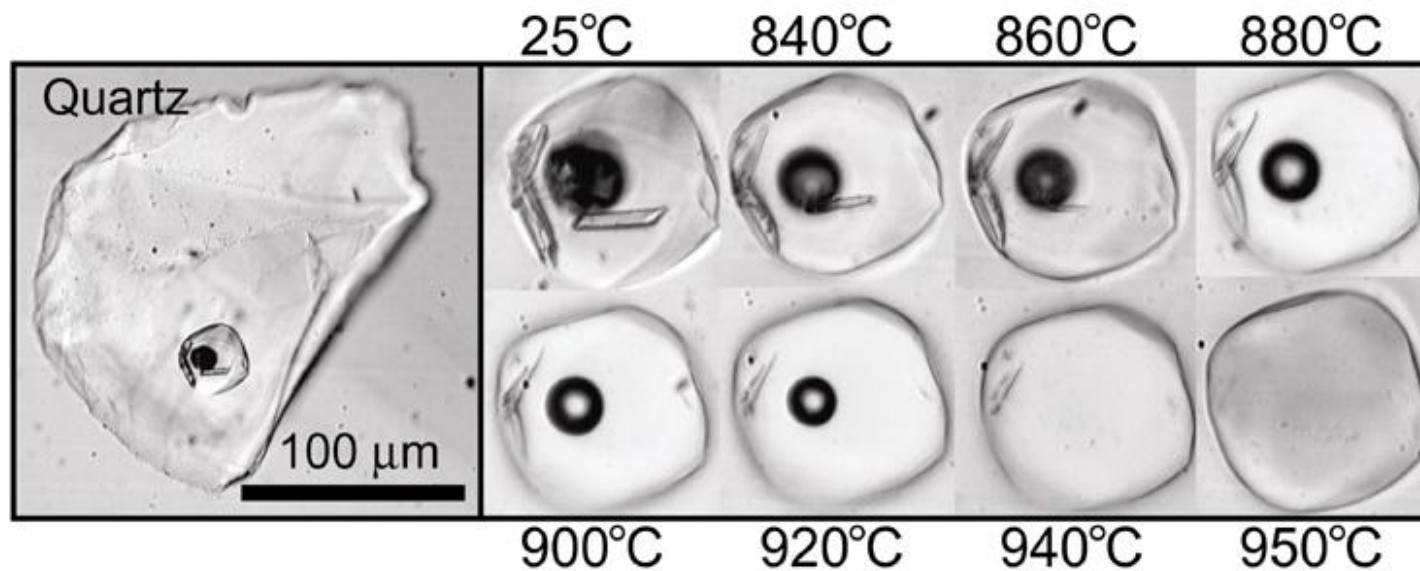
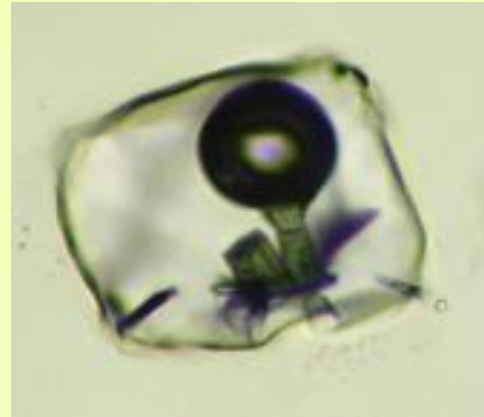
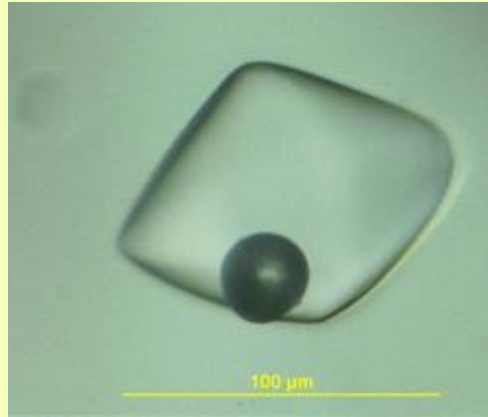
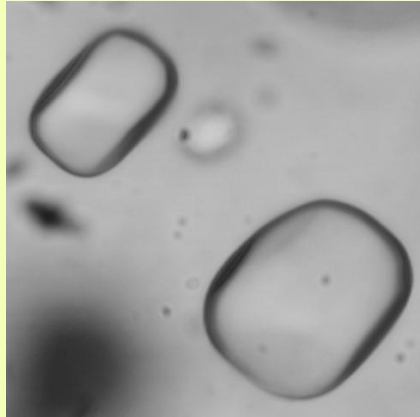


When the inclusion is heated, the small film of H<sub>2</sub>O that occurs wetting the walls at room temperature evaporates into the CO<sub>2</sub>-rich fluid phase to produce a homogeneous H<sub>2</sub>O-CO<sub>2</sub> fluid. Evaporation of the H<sub>2</sub>O into the CO<sub>2</sub> produces a larger analytical volume containing H<sub>2</sub>O, and the H<sub>2</sub>O peak becomes sharper and more intense with increasing temperature.

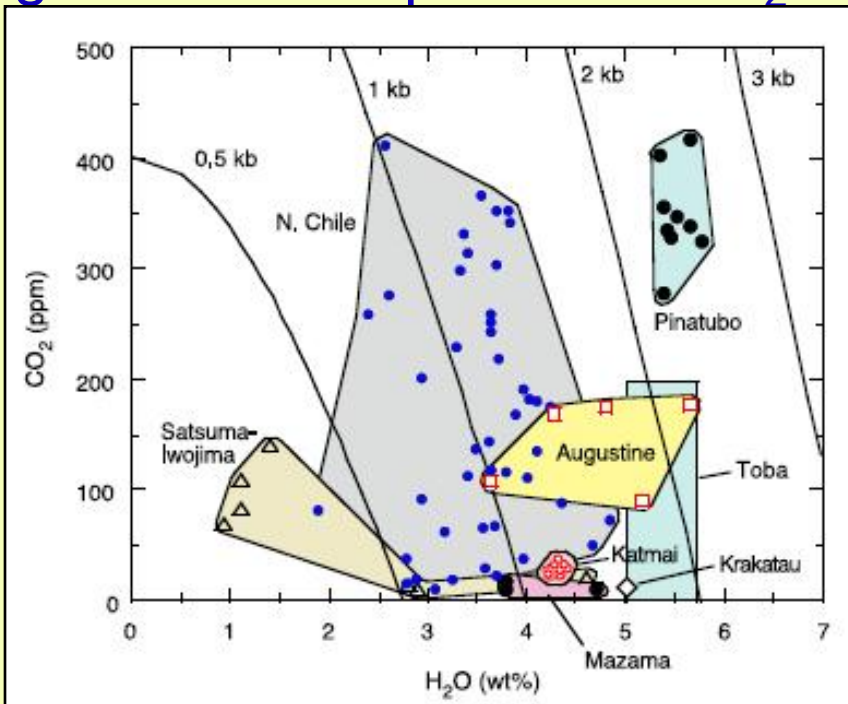
In some fluid inclusions from granulites, a small amount of  $\text{CH}_4$ , in addition to the more abundant  $\text{CO}_2$ , is also detected.



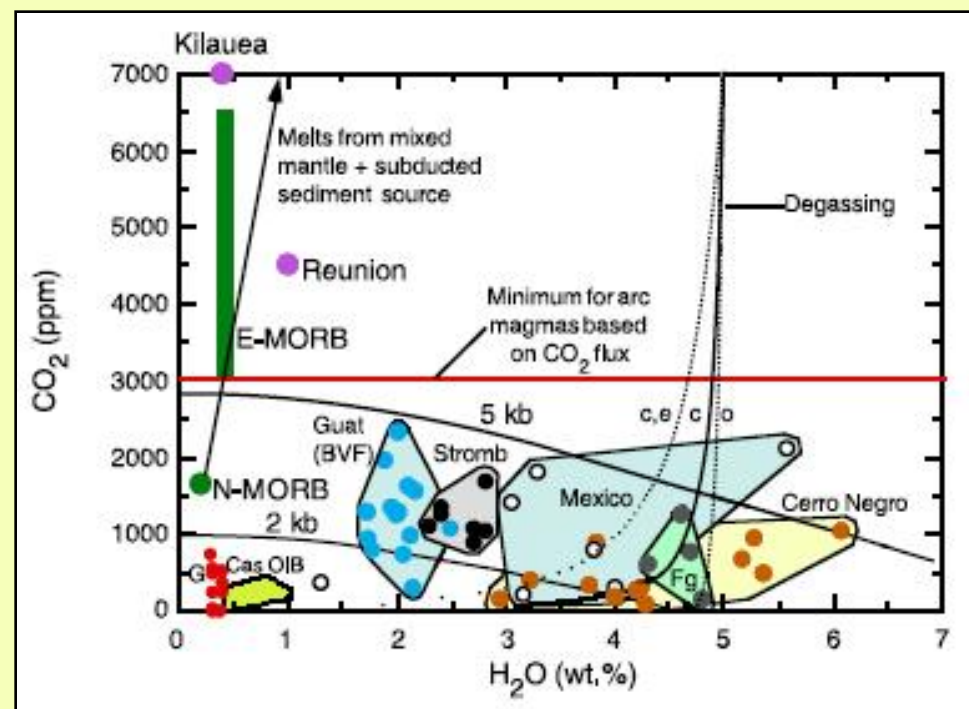
Melt inclusions trapped in phenocrysts in magmatic systems provide the best and most direct evidence for volatile abundances and speciation in lithospheric magmas.



Carbon dioxide contents of melt inclusions from arc basalts (MI mostly in olivine) are significantly higher than CO<sub>2</sub> content of MI from arc dacites and rhyolites (MI mostly in quartz). The differences in CO<sub>2</sub> content largely reflect differences in magma generation depths and CO<sub>2</sub> solubilities in mafic vs. felsic melts.

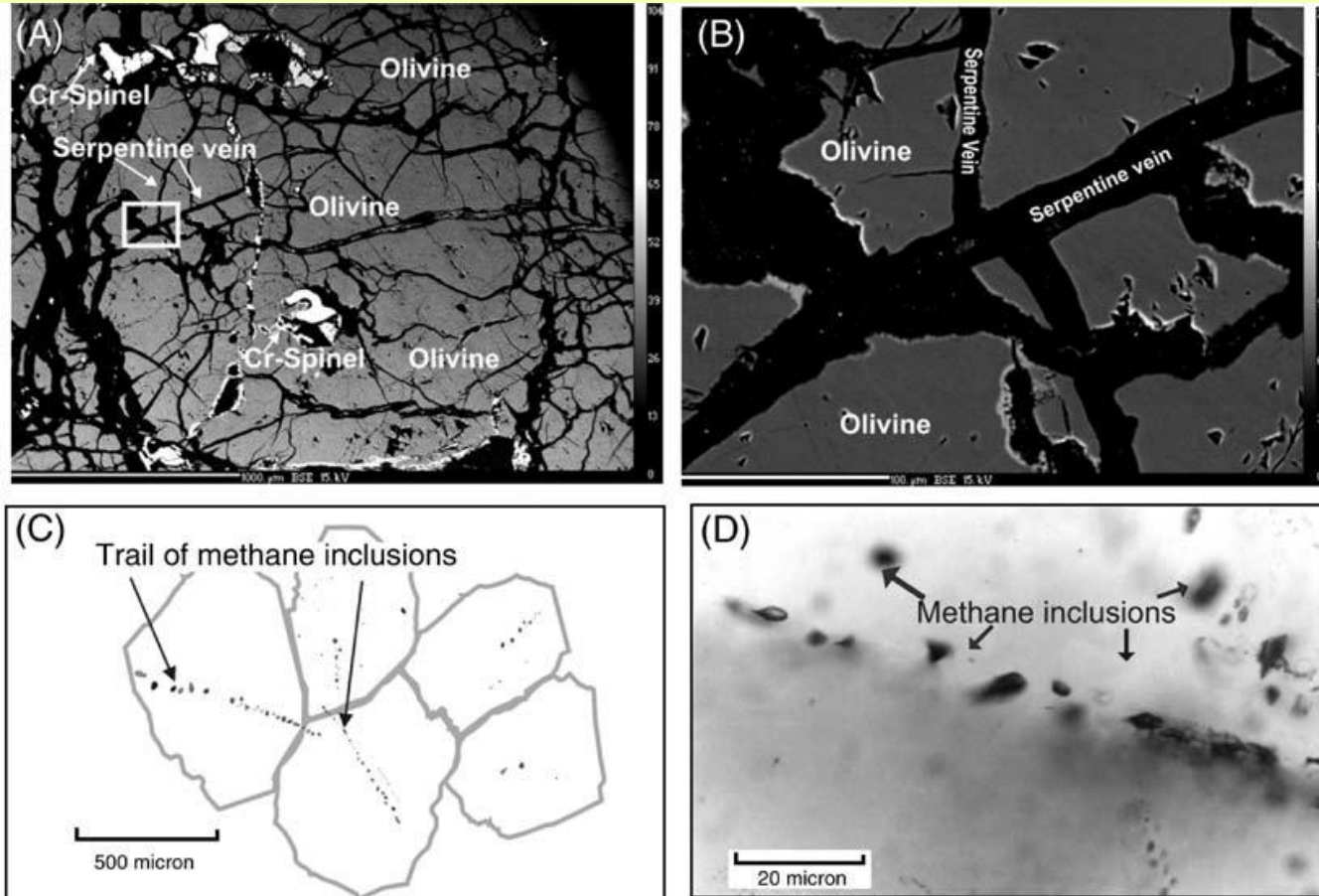


Arc dacites and rhyolites



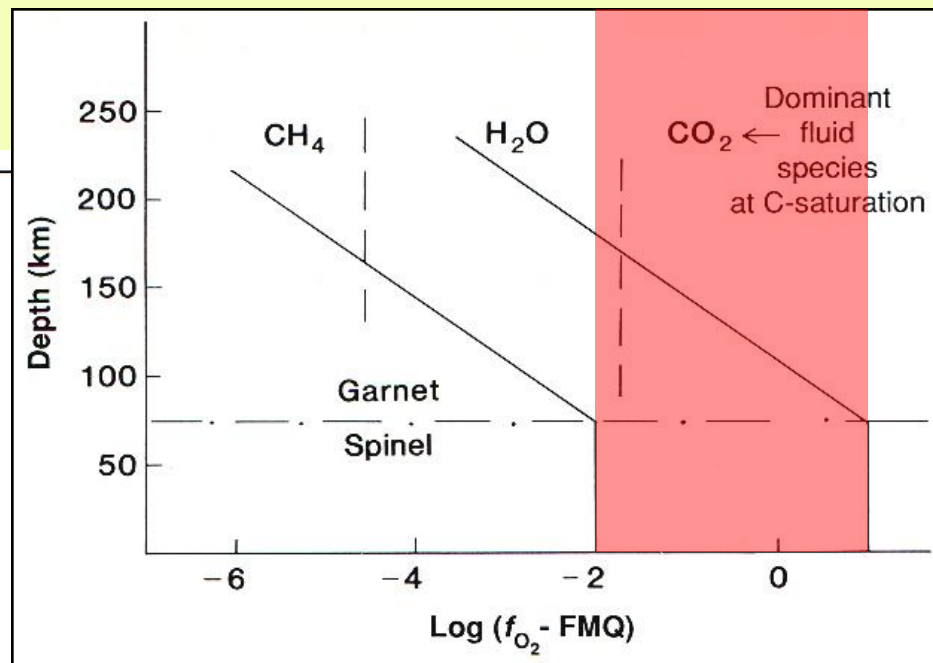
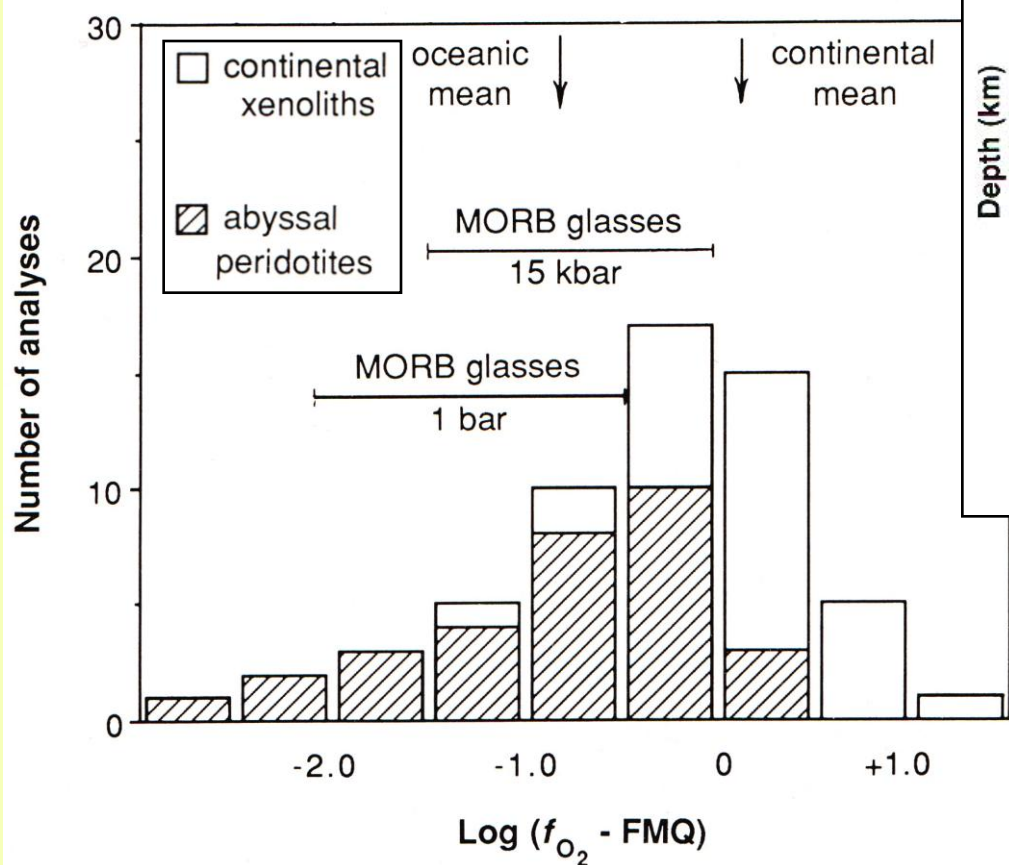
Arc basalts

In submarine hydrothermal systems  $\text{CH}_4$  is generated during the serpentinization of olivine. Sometimes  $\text{CH}_4$  produced during serpentinization is preserved in secondary inclusions in rocks that have experienced rapid burial and uplift.

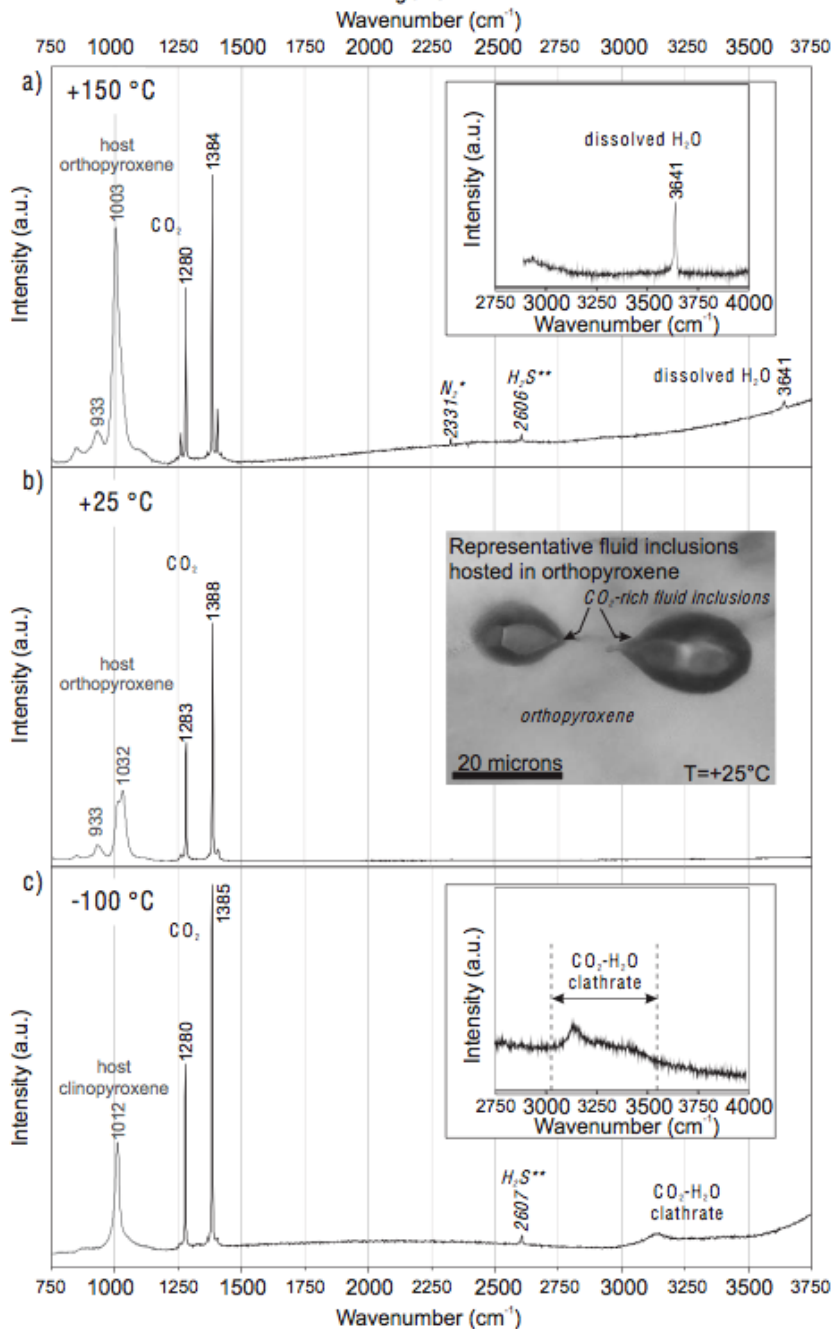


From Sachan et. al., EPSL, 2007

Mantle xenoliths indicate that the oxygen fugacity in the upper mantle is within one order of magnitude of QFM and that at mantle PT conditions  $\text{CO}_2 \pm \text{H}_2\text{O}$  should be the stable fluid species.

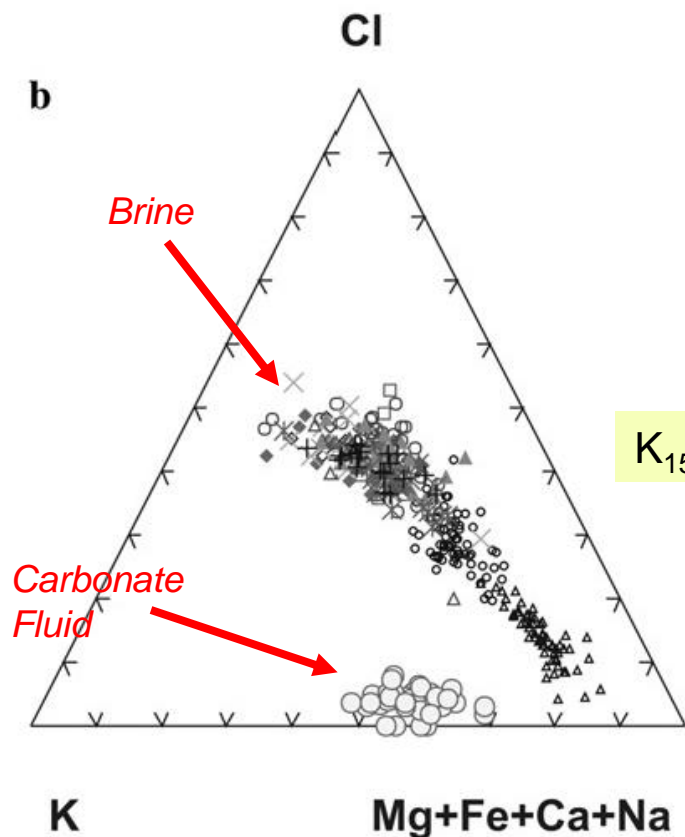


Diagonal lines on the figure above show how the oxygen fugacity should vary in the garnet field if the peridotites contain the same amount of  $\text{Fe}^{3+}$  as in the spinel field.



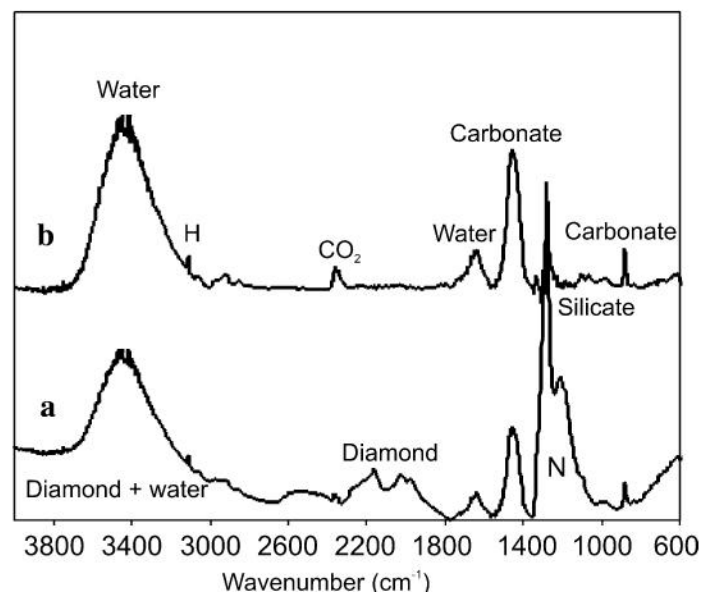
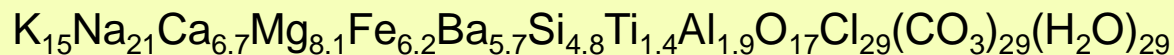
Peridotitic mantle xenoliths often appear contain “pure” CO<sub>2</sub> fluid inclusions. However, when analyzed by Raman spectroscopy at either low (-100° C) or elevated (+150° C) temperature, a distinct peak for H<sub>2</sub>O is observed. The H<sub>2</sub>O peak is not present when the same inclusion is analyzed at room temperature. Work in progress suggests that the presence of small amounts of H<sub>2</sub>O in CO<sub>2</sub>-rich FI from upper mantle environments is common.

# Analyses of fluid inclusions in diamond suggests the presence of carbonate-rich fluids and CO<sub>2</sub> in the source region.



From Klein-BenDavid et al., GCA 71 (2007)

Diamond from the Diavik Mine, Canada, contains brine and carbonate inclusions. The two fluids are thought to represent immiscibility of an original high-density carbonatitic melt to produce one fluid rich in Si, H<sub>2</sub>O and Cl, and another carbonate-rich fluid. The carbonate end-member composition is approximately:

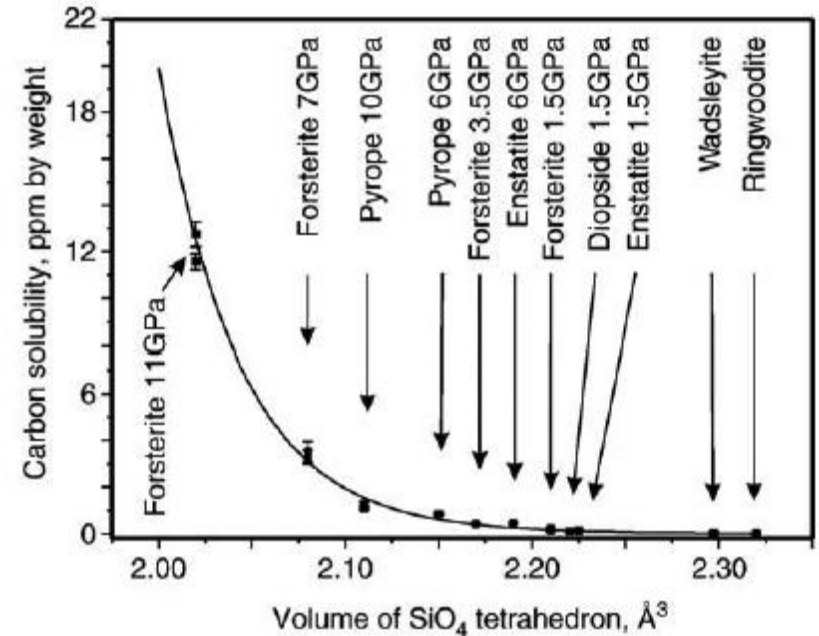
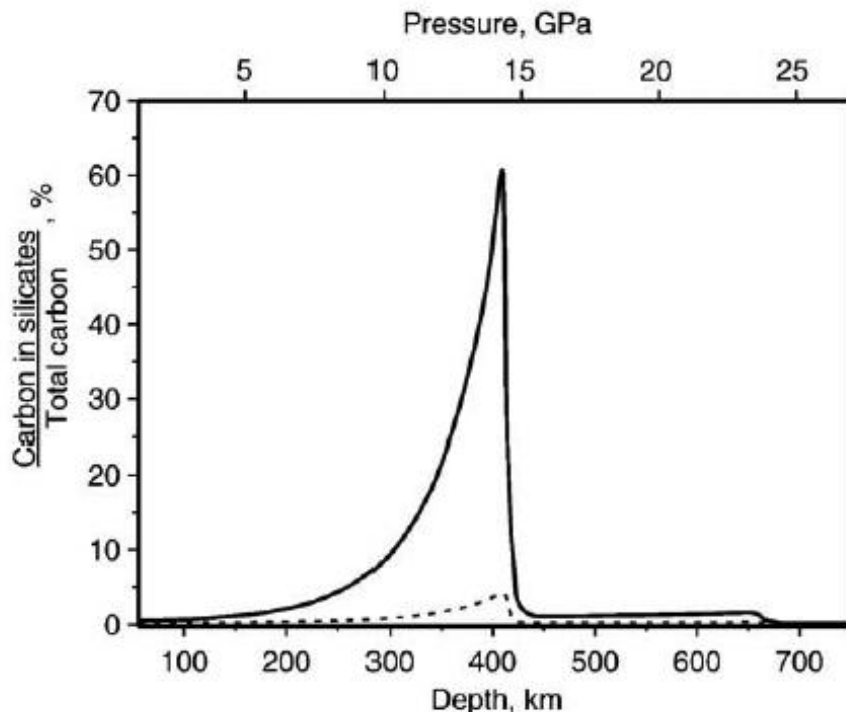


*FTIR spectrum of inclusions in diamond.*

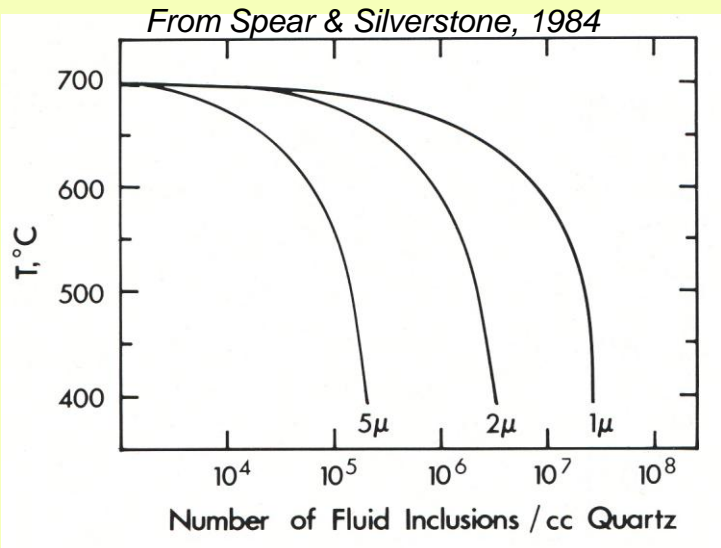
*(Top) After subtraction of the diamond and nitrogen peaks.*

Recent work suggests that in the deeper mantle most carbon is present as defects (nano- to pico-scale FI?) in volatile-free minerals.

Carbon dissolved in volatile-free minerals can exsolve as the minerals are exposed to lower pressure conditions as a result of mantle convection.

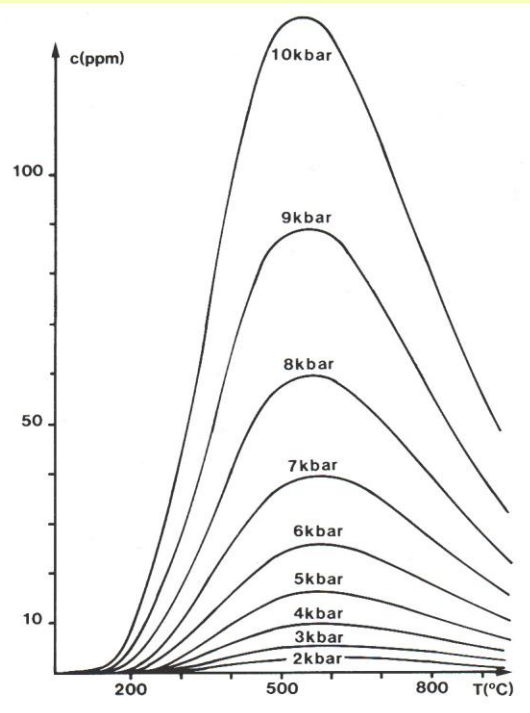
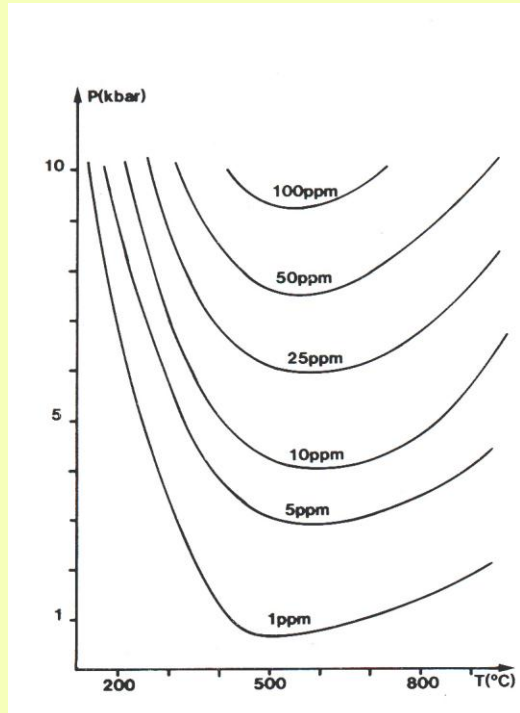


The highest proportion of carbon dissolved in volatile-free minerals is expected to occur in a thin zone of the upper mantle just above the transition zone.



If water is dissolved in quartz at high temperature and pressure as an OH defect, how much fluid can be generated if the  $\text{H}_2\text{O}$  exsolves during uplift?

$\approx 10^7 \text{ kg H}_2\text{O per km}^3 \text{ rock!}$



Several tens of ppm  $\text{H}_2\text{O}$  can dissolve in quartz at high temperature and pressure. When this quartz is uplifted to the surface water may exsolve from the quartz structure to produce tiny, uniformly distributed fluid inclusions that give quartz a milky texture.

From Doukhan & Paterson,

# Summary

- Fluid and melt inclusions provide the most direct evidence for the composition and distribution of carbon in the Earth.
- Carbon speciation is controlled by the oxidation state of the system, which in turn depends on solid phases present.
- In organic-rich sedimentary environments and the anchi-metamorphic zone the dominant carbon species is  $\text{CH}_4$ .
- In metamorphic environments fluids evolve from being  $\text{CH}_4$ -rich to  $\text{CO}_2$ -rich with increasing metamorphic grade.
- Silicate melts contain from a few 100 (felsic) to a few 1000 (mafic) ppm  $\text{CO}_2$ .
- In the deeper upper mantle most carbon is likely present as nano- to pico-scale FI (defects) in volatile-free minerals. The dissolved C can exsolve during mantle convection to produce  $\text{CO}_2$ -rich FI that are ubiquitous in upper mantle xenoliths.